

# Classical Statistical Mechanics: Part 2

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## Variational approach to equilibrium densities

- Hypothesis: functional form of equilibrium distribution is that which maximizes the statistical entropy.
  - Statistical entropy viewed as a functional of  $f(\mathbf{x}^{(N)})$ .
  - Form must satisfy basic properties of a probability density: normalized.

## Microcanonical Ensemble

- All systems in ensemble have an energy in the energy shell  $[E, E + \delta E]$ .
- Task is to maximize

$$S = -k \int_{E < H(\mathbf{x}^{(N)}) < E + \delta E} d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) \ln (h^{3N} N! f(\mathbf{x}^{(N)})).$$

subject to the constraint

$$\int_{E < H(\mathbf{x}^{(N)}) < E + \delta E} d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) \equiv \int' d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) = 1.$$

- Procedure:

$$\begin{aligned} \delta(S + \alpha_0 1) &= \int' d\mathbf{x}^{(N)} [-k f(\mathbf{x}^{(N)}) \ln (N! h^{3N} f(\mathbf{x}^{(N)})) + \alpha_0 f(\mathbf{x}^{(N)})] = 0 \\ &= \int' d\mathbf{x}^{(N)} [-k \ln (N! h^{3N} f(\mathbf{x}^{(N)})) - k + \alpha_0] \delta f(\mathbf{x}^{(N)}) = 0. \end{aligned}$$

– Since the variation is arbitrary:

$$-k \ln (N!h^{3N} f(\mathbf{x}^{(N)})) - k + \alpha_0 = 0$$

so

$$\begin{aligned} k \ln (N!h^{3N} f(\mathbf{x}^{(N)})) &= \alpha_0 - k \\ N!h^{3N} f(\mathbf{x}^{(N)}) &= e^{\alpha_0/k-1} = \tilde{C}. \end{aligned}$$

– Hence, as desired, we get:

$$f(\mathbf{x}^{(N)}) = \begin{cases} C & \text{for } E < H(\mathbf{x}^{(N)}) < E + \delta E \\ 0 & \text{otherwise} \end{cases}$$

where  $C$  is a constant.

– From normalization,  $C = 1/\Omega$ .

- It is cumbersome to work with integrals over restricted energy shells. Can we relax this restriction and obtain ensemble averages that agree with time averages (and each other)?

### Canonical Ensemble

- Remove restriction of defining probability only on constant energy hypersurface.
- Allow total energy of systems in ensemble to vary (hopefully) narrowly around a fixed average value.
- Task now is to maximize

$$S = -k \int d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) \ln (h^{3N} N! f(\mathbf{x}^{(N)})).$$

subject to the constraints

$$\begin{aligned} \int d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) &= 1 \\ \bar{E} &= \int d\mathbf{x}^{(N)} H(\mathbf{x}^{(N)}) f(\mathbf{x}^{(N)}) \end{aligned}$$

- Procedure:

$$\delta \int d\mathbf{x}^{(N)} \left[ -k f(\mathbf{x}^{(N)}) \ln (N! h^{3N} f(\mathbf{x}^{(N)})) + \alpha_0 f(\mathbf{x}^{(N)}) + \alpha_E H(\mathbf{x}^{(N)}) f(\mathbf{x}^{(N)}) \right] = 0$$

So

$$\alpha_0 + \alpha_E H(\mathbf{x}^{(N)}) - k \ln (N! h^{3N} f(\mathbf{x}^{(N)})) - k = 0$$

$$f(\mathbf{x}^{(N)}) = \frac{1}{N! h^{3N}} \exp \left\{ \frac{\alpha_0}{k} - 1 + \frac{\alpha_E}{k} H(\mathbf{x}^{(N)}) \right\}$$

- What are the Lagrange multipliers?

1. From normalization:

$$\int d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) = 1 = \frac{\exp\{\alpha_0/k - 1\}}{N! h^{3N}} \int d\mathbf{x}^{(N)} \exp \left\{ \frac{\alpha_E}{k} H(\mathbf{x}^{(N)}) \right\}$$

$$\exp\{1 - \alpha_0/k\} = \frac{1}{N! h^{3N}} \int d\mathbf{x}^{(N)} \exp \left\{ \frac{\alpha_E}{k} H(\mathbf{x}^{(N)}) \right\}$$

2. From the thermodynamic relation:

$$\left( \frac{\partial S}{\partial \bar{E}} \right)_{v, \text{norm}} = \frac{1}{T}$$

- Note from above that since  $k \ln(N! h^{3N} f(\mathbf{x}^{(N)})) = \alpha_E H(\mathbf{x}^{(N)}) + \alpha_0 - k$ , we have

$$S = - \int d\mathbf{x}^{(N)} f(\mathbf{x}^{(N)}) (\alpha_E H(\mathbf{x}^{(N)}) + \alpha_0 - k)$$

$$= (k - \alpha_0) - \alpha_E \bar{E}.$$

- It is now clear that:

$$\left( \frac{\partial S}{\partial \bar{E}} \right)_{v, \text{norm}} = \frac{1}{T} = -\alpha_E$$

$$\alpha_E = -\frac{1}{T}.$$

- Thus  $TS = T(k - \alpha_0) + \bar{E}$  or  $\bar{E} - TS = T(\alpha_0 - k)$ .

–  $\bar{E} - TS = A$ , the Helmholtz free energy, so

$$T(\alpha_0 - k) = A$$

and  $\exp\{\alpha_0/k - 1\} = \exp\{\beta A\}$  where  $\beta = 1/(kT)$ .

- We conclude that for the canonical ensemble:

$$f(\mathbf{x}^{(N)}) = \frac{1}{N!h^{3N}} \exp\{\beta(A - H(\mathbf{x}^{(N)}))\}$$

- We define the *partition function*  $Q_N(T, V)$  by

$$Q_N(T, V) = \frac{1}{N!h^{3N}} \int d\mathbf{x}^{(N)} \exp\{-\beta H(\mathbf{x}^{(N)})\} = \exp\{-\beta A\}$$

so

$$f(\mathbf{x}^{(N)}) = \frac{1}{N!h^{3N}} \exp\{\beta(A - H(\mathbf{x}^{(N)}))\} = \frac{1}{N!h^{3N}} \frac{\exp\{-\beta H(\mathbf{x}^{(N)})\}}{Q_N(T, V)}.$$

- Relation  $A = -kT \ln Q_N(T, V)$  gives thermodynamic connection: For example

1. The pressure is:

$$P = - \left( \frac{\partial A}{\partial V} \right)_T = kT \left( \frac{\partial \ln Q_N}{\partial V} \right)_T.$$

2. The chemical potential is:

$$\mu = \left( \frac{\partial A}{\partial N} \right)_{T, V}$$

3. The energy is:

$$\begin{aligned} \bar{E} &= \frac{\exp\{\beta A\}}{N!h^{3N}} \int d\mathbf{x}^{(N)} H(\mathbf{x}^{(N)}) \exp\{-\beta H(\mathbf{x}^{(N)})\} \\ &= \frac{\exp\{\beta A\}}{N!h^{3N}} - \frac{\partial}{\partial \beta} \int d\mathbf{x}^{(N)} \exp\{-\beta H(\mathbf{x}^{(N)})\} \\ &= -\frac{1}{Q_N} \frac{\partial Q_N}{\partial \beta} = -\frac{\partial \ln Q_N}{\partial \beta}. \end{aligned}$$

- We have obtained two different probability densities corresponding to different “ensembles”.
  - Microcanonical ensemble: vanishing probability to find a system in ensemble with energy function (Hamiltonian) different from average energy.
  - Canonical ensemble: Hamiltonian can differ among states but states must be weighted so that the ensemble average of Hamiltonian corresponds to the average energy.
- What is the correspondence between averages in the different ensembles?
- We can write the canonical partition function as:

$$\begin{aligned}
 Q_N(T, V) &= \frac{1}{N!h^{3N}} \int d\mathbf{x}^{(N)} \exp\{-\beta H(\mathbf{x}^{(N)})\} \\
 &= \int_0^\infty dE \frac{1}{N!h^{3N}} \int d\mathbf{x}^{(N)} \exp\{-\beta H(\mathbf{x}^{(N)})\} \delta(E - H(\mathbf{x}^{(N)})) \\
 &= \int_0^\infty dE \exp\{-\beta E\} \left( \frac{1}{N!h^{3N}} \int d\mathbf{x}^{(N)} \delta E - H(\mathbf{x}^{(N)}) \right) \\
 Q_N(T, V) &= \int_0^\infty dE \exp\{-\beta E\} N(E)
 \end{aligned}$$

where

$$\begin{aligned}
 N(E) &\equiv \frac{\tilde{N}(E)}{N!h^{3N}} \\
 &= \text{density of } \textit{unique} \text{ states at energy } E \text{ (microcanonical partition function)}.
 \end{aligned}$$

- Partition function  $Q_N(\beta, V)$  is the Laplace transform of density of unique states.
- $Q_N$  is like a generating function for  $N(E)$ .
- Recall the definition of the statistical entropy:

$$\begin{aligned}
 S(E) &= k \ln W = k \ln \left( \frac{\Omega(E, \delta E, N, V)}{N!h^{3N}} \right) \\
 &\approx k \ln(N(E)\delta E) = k \ln N(E) + C
 \end{aligned}$$

for small  $\delta E$  where  $\Omega \approx \tilde{N}(E)\delta E$ .

- Thus

$$N(E) \sim \exp\{S(E)/k\} \quad Q_N(T, V) \sim \int_0^\infty dE \exp\{-\beta(E - TS)\}$$

- Typically,  $N(E)$  is a rapidly increasing function of  $E$

$$N(E) \sim f(E/N, V/N) \exp\{Ng(E/n, V/N)\} \quad S(E) \sim Nk g(E/N, V/N).$$

### Thermodynamic Stability

- Define  $P(E)dE$  to be the probability of finding a system with energy  $E$  in the canonical ensemble.

$$P(E) = \langle \delta(E - H(\mathbf{x}^{(N)})) \rangle = \int d\mathbf{x}^{(N)} \delta(E - H(\mathbf{x}^{(N)})) f(\mathbf{x}^{(N)}) = \exp\{\beta A\} \exp\{-\beta E\} N(E)$$

- From normalization, it follows that

$$\exp\{-\beta A\} = \int_0^\infty dE \exp\{-\beta E\} N(E).$$

- At low temperatures, how much more likely are we to find a system with  $E = E_1$  than  $E = E_2$ ?

$$\begin{aligned} \frac{P(E_1)}{P(E_2)} &= \frac{e^{\beta A} e^{-\beta E_1} N(E_1)}{e^{\beta A} e^{-\beta E_2} N(E_2)} \\ &= e^{-\beta \Delta E} \frac{N(E_1)}{N(E_2)} \quad \Delta E \equiv E_1 - E_2 \end{aligned}$$

- Since  $S(E) \sim k \ln(N(E))$ ,

$$\frac{P(E_1)}{P(E_2)} = e^{-\beta(\Delta E - T\Delta S)} \quad \Delta S \equiv S(E_1) - S(E_2)$$

- The density of states is important (entropic factor).
- Probabilities determined by balance between  $\Delta E$  and  $\Delta S$ .

- Consider a bio-molecule (like a polymer chain): Each phase point corresponds to a particular configuration (and set of momenta) of the molecule.

- Suppose there are two different *types* of configuration  $\Gamma_1$  and  $\Gamma_2$ : represented by *basins* in configurational space.
- Define conditional probabilities partitioning configurational space:

$$P(E) = \sum_i P(E|\Gamma_i)$$

$$\begin{aligned} P(E|\Gamma_1) &= \int_{\Gamma_1} d\mathbf{x}^{(N)} \delta(E - H(\mathbf{x}^{(N)})) f(\mathbf{x}^{(N)}) = \frac{e^{\beta A} e^{-\beta E}}{N! h^{3N}} \int_{\Gamma_1} d\mathbf{x}^{(N)} \delta(E - H(\mathbf{x}^{(N)})) \\ &= e^{\beta A} e^{-\beta E} N(E|\Gamma_1) \\ N(E|\Gamma_1) &= \int_{\Gamma_1} d\mathbf{x}^{(N)} \frac{\delta(E - H(\mathbf{x}^{(N)}))}{N! h^{3N}} = \# \text{ of unique states in basin } \Gamma_1. \end{aligned}$$

- Thus

$$\frac{P(E_1|\Gamma_1)}{P(E_2|\Gamma_2)} = e^{-\beta \Delta E} \frac{N(E_1|\Gamma_1)}{N(E_2|\Gamma_2)}$$

gives the relative stability of the conformations at the given energies.

### Relationship between ensemble averages

- How likely are we to observe a system in the canonical ensemble with an energy very different from the average energy  $\bar{E} = \langle H(\mathbf{x}^{(N)}) \rangle$ ?

$$Pr(|H(\mathbf{x}^{(N)}) - \bar{E}| \geq \lambda \bar{E}) \leq \frac{\sigma_E^2}{\lambda^2 \bar{E}^2}$$

- Now the variance in the energy is:

$$\sigma_E^2 = \langle H(\mathbf{x}^{(N)})^2 \rangle - \langle H(\mathbf{x}^{(N)}) \rangle^2 = \frac{\partial^2 \ln Q_N}{\partial \beta^2} = -\frac{\partial \bar{E}}{\partial \beta} = kT^2 C_v$$

and hence

$$Pr(|H(\mathbf{x}^{(N)}) - \bar{E}| \geq \lambda \bar{E}) \leq \frac{kT^2 C_v}{\lambda^2 \bar{E}^2}$$

- We will show that for an ideal gas system,  $\bar{E} = 3/2NkT$  and hence  $C_v = 3/2Nk$ .
- Typically,  $\bar{E} \sim N$  and  $C_v \sim N$ .

$$Pr(|H(\mathbf{x}^{(N)}) - \bar{E}| \geq \lambda \bar{E}) \leq \frac{kT^2 C_v}{\lambda^2 \bar{E}^2} \sim \frac{1}{N\lambda^2}$$

- As  $N$  increases, it becomes less and less likely to observe a system with energy very different from  $\bar{E}$ ,

$$\langle B(\mathbf{x}^{(N)}) \rangle_{\text{canon}} = \int dE P(E) \langle B(\mathbf{x}^{(N)}) \rangle_{\text{micro at } E} \approx \langle B(\mathbf{x}^{(N)}) \rangle_{\text{micro at } \bar{E}} (1 + O(1/N)).$$

- $P(E)$  is sharply-peaked around  $E = \bar{E}$ : Can show (homework?)

$$P(E) \approx P(\bar{E}) \left( \frac{1}{2\pi\sigma_E^2} \right)^{1/2} \exp \left\{ -\frac{(E - \bar{E})^2}{2kT^2 C_v} \right\}$$

- Relative spread of energy  $\sigma_E/\bar{E} \sim N^{-1/2}$ .